



1. Vectors

Force, velocity and acceleration which involve both a magnitude and direction, are **vectors**. A vector is written using a bold typeface, \mathbf{a} , or an underline \underline{a} . It is represented pictorially by a **directed line segment** as shown.

The length of the line segment represents the vector's magnitude. Its orientation, together with the arrow shown, gives the direction of the vector. The magnitude of a vector \underline{a} is written $|\underline{a}|$ or simply a . A **unit vector** has magnitude 1. $-\underline{a}$ has the magnitude of \underline{a} but is opposite in direction.

Addition: The parallelogram rule defines addition of two vectors: $\underline{c} = \underline{a} + \underline{b}$ where θ is the angle between \underline{a} and \underline{b} , as shown. \underline{c} is called the **resultant** of \underline{a} and \underline{b} .

Rectangular Components: Let \underline{i} be a unit vector in the direction of the positive x axis and \underline{j} be a unit vector in the direction of the positive y axis. In two dimensions the vector \underline{a} can be written as the sum of two rectangular vector components: $\underline{a} = a_1 \underline{i} + a_2 \underline{j}$ or $\underline{a} = (a_1, a_2)$. The scalar components a_1 and a_2 are given by $a_1 = a \cos \theta$, $a_2 = a \sin \theta$, where θ is the angle \underline{a} makes with the positive x axis. Any vector can be replaced by its rectangular vector components, starting at the same point.

Using Pythagoras' theorem it follows that $|\underline{a}| = \sqrt{a_1^2 + a_2^2}$. In a natural extension to three dimensions we can write $|\underline{a}| = \sqrt{a_1^2 + a_2^2 + a_3^2}$ and $\underline{a} = a_1 \underline{i} + a_2 \underline{j} + a_3 \underline{k}$.

Scalar (dot) product & Vector (cross) product: If $\underline{a} = a_1 \underline{i} + a_2 \underline{j} + a_3 \underline{k}$ and $\underline{b} = b_1 \underline{i} + b_2 \underline{j} + b_3 \underline{k}$ then $\underline{a} \cdot \underline{b} = a_1 b_1 + a_2 b_2 + a_3 b_3$ and $|\underline{a}| |\underline{b}| \cos \theta = \underline{a} \cdot \underline{b}$ where θ is the angle between \underline{a} and \underline{b} . Here θ is the angle between \underline{a} and \underline{b} , and \underline{n} is a unit vector perpendicular to the plane containing \underline{a} and \underline{b} in a sense defined by the right-hand screw rule.

$$\underline{a} \times \underline{b} = \begin{vmatrix} \underline{i} & \underline{j} & \underline{k} \\ a_1 & a_2 & a_3 \\ b_1 & b_2 & b_3 \end{vmatrix} = (a_2 b_3 - a_3 b_2) \underline{i} + (a_3 b_1 - a_1 b_3) \underline{j} + (a_1 b_2 - a_2 b_1) \underline{k}$$

Quantity	Unit	Symbol
Mass	kilogram	kg
Length	metre	m
Time	second	s
Force	newton	N (1 N = 1 kg m s ⁻²)
Work	joule	J (1 J = 1 N m)
Power	watt	W (1 W = 1 J s ⁻¹)
Velocity	metre per second	m s ⁻¹
Acceleration	metre per second squared	m s ⁻²
Energy	joule	J
Momentum	newton second	N s
Impulse	newton second	N s
Angular Velocity/Angular Frequency	radians per second	rad s ⁻¹

The SI system uses the following units:

3. Units

$G = 6.673 \times 10^{-11} \text{ m}^3 \text{ kg}^{-1} \text{ s}^{-2}$.

is called the gravitational constant. Its accepted value is m_2 are their masses, r is the distance between them. G is the magnitude of the gravitational force on either body, m_1 and m_2 are their masses, r is the distance between them. Thus $F_g = G \frac{m_1 m_2}{r^2}$ where F_g is the magnitude of the gravitational force on either body, m_1 and m_2 are their masses, r is the distance between them. Thus $F_g = G \frac{m_1 m_2}{r^2}$ where F_g is the magnitude of the gravitational force on either body, m_1 and m_2 are their masses, r is the distance between them.

Newton's Law of Universal Gravitation: Every body in the universe attracts every other body with a force which is directly proportional to the product of the masses and inversely proportional to the square of the distance between them. Thus $F_g = G \frac{m_1 m_2}{r^2}$ where F_g is the magnitude of the gravitational force on either body, m_1 and m_2 are their masses, r is the distance between them.

When bodies interact. Whenever body A exerts a force, \underline{F}_A , of magnitude F_A on body B, B exerts a force, $-\underline{F}_B$, on body A.

To every action there is an equal and opposite reaction. Thus forces come in pairs positive reaction. Whenever body A exerts a force, \underline{F}_A , of magnitude F_A on body B, B exerts a force, $-\underline{F}_B$, on body A.

Newton's third law of motion:

and $\underline{a} = (a_x, a_y, a_z)$.

$\underline{F} = m \underline{a}$, $F_x = ma_x$, $F_y = ma_y$, $F_z = ma_z$ where $\underline{F} = (F_x, F_y, F_z)$ and $\underline{a} = (a_x, a_y, a_z)$.

Newton's second law of motion: If a body of mass m is moving with velocity \underline{v} , and so has momentum $m\underline{v}$, then the rate of change of momentum of the body is directly proportional to the resultant applied force, \underline{F} , acting on it: $\underline{F} = \frac{d(m\underline{v})}{dt}$. For a body with acceleration \underline{a} and of constant mass m , this becomes $\underline{F} = m \frac{d\underline{v}}{dt} = m\underline{a}$. This vector equation is equivalent to the scalar equations: $F_x = ma_x$, $F_y = ma_y$, $F_z = ma_z$ where $\underline{F} = (F_x, F_y, F_z)$ and $\underline{a} = (a_x, a_y, a_z)$.

Newton's first law of motion: A body will remain at rest or continue its uniform motion in a straight line unless compelled to change by forces acting on it. It follows from this that when a body is in equilibrium the resultant force, $\underline{R} = (R_x, R_y, R_z)$, of all the forces acting on it, is zero. Thus $R_x = 0$, $R_y = 0$, $R_z = 0$ where R_x, R_y and R_z are the net sums of the x, y and z scalar components of the forces, respectively.

2. Newton's Laws of Motion and Gravitation

13. Rigid bodies

Consider an axis perpendicular to the plane of the paper and passing through O. The rigid body is acted upon by the forces \underline{F}_1 and \underline{F}_2 , lying in the plane. $\underline{F}_1, \underline{F}_2$ produce anti-clockwise/clockwise rotation about the axis, respectively. By convention, anti-clockwise rotation is taken as positive. The **moments** of \underline{F}_1 and \underline{F}_2 about the axis through O are defined by

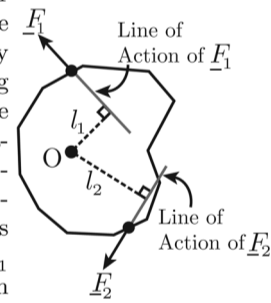
$$\Gamma_1 = +F_1 l_1 \quad \Gamma_2 = -F_2 l_2$$

where l_1 and l_2 are the perpendicular distances of the lines of action of \underline{F}_1 and \underline{F}_2 from O. The line of action of a force is a line with the same orientation as the force and which passes through its point of action.

For rigid bodies there are two necessary conditions for equilibrium:

First condition: When a body is in equilibrium the resultant force, $\underline{R} = (R_x, R_y, R_z)$, of all the forces acting on it, is zero. (This condition also applies to particles.) Thus $\underline{R} = \underline{0}$ $R_x = 0$ $R_y = 0$ $R_z = 0$ where R_x, R_y and R_z are the net sums of the x, y and z scalar components of the forces, respectively.

Second condition: When a body is in equilibrium the sum of the moments, about any arbitrary axis, is zero: $\Sigma \Gamma = 0$



Centre of mass: This is the point in a body such that an external force produces an acceleration just as though the whole mass were concentrated there. Let $(\bar{x}, \bar{y}, \bar{z})$ be the coordinates of the centre of mass of a system of particles, each of mass m_1, m_2, \dots , and centres of mass located at $(x_1, y_1, z_1), (x_2, y_2, z_2), \dots$. Then

$$\bar{x} = \frac{\Sigma m_i x_i}{\Sigma m_i} \quad \bar{y} = \frac{\Sigma m_i y_i}{\Sigma m_i} \quad \bar{z} = \frac{\Sigma m_i z_i}{\Sigma m_i}$$

from which $\Sigma m_i (x_i - \bar{x}) = \Sigma m_i (y_i - \bar{y}) = \Sigma m_i (z_i - \bar{z}) = 0$

Then the sum of moments about an axis through the centre of mass is zero. Symmetry can be useful in finding the centre of mass. The centre of mass of a homogeneous sphere, circular disk or rectangular plate is at its centre.

Written by Dr. Carol Robinson¹, Dr. Tony Croft¹, & Prof. Mike Savage² with additional comments by Dr. Marie Bassford³ Images produced by Paul Newman¹

¹ Mathematics Education Centre
² Department of Physics & Astronomy University of Leeds
³ Engineering Education Centre ©Loughborough University 2003

$T = \frac{F}{\lambda}$ where λ is Young's modulus of elasticity, or $T = kx$ where $k = \frac{T}{\lambda}$ is called the spring stiffness.

Hooke showed that, provided the extension is not too great, the tension, T , in an elastic string or spring is directly proportional to the extension, x , and inversely proportional to its natural length, L .

Tension: (iii) Elastic strings or springs (Hooke's Law). Hooke showed that, provided the extension is not too great, the tension, T , in an elastic string or spring is directly proportional to the extension, x , and inversely proportional to its natural length, L .

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Tension: (i) Light, inextensible strings. A mass m hangs in equilibrium on the end of an inextensible string attached to a ceiling, diagram (a). The tension at any point of the string equals the force exerted at that point. The string is said to be 'light' if its weight is negligible compared to the weight mg , and the tension, T , is then constant along its length. A separated body diagram (b) shows the forces acting on the mass and the string, and shows the tension exerted on the ceiling. From Newton's 2nd Law $T = mg$.

Reaction: A block, of mass m , rests on a horizontal surface, as shown in diagram (a). The block and the surface interact, exerting on each other equal and opposite normal reactions of magnitude R . A separated body diagram for the block is shown in diagram (b). Since the block is at rest $R = mg$ from Newton's 2nd Law.

Weight: The weight of a body, of mass m , is defined to be the force, \underline{W} , which it is attracted to the Earth. Its magnitude, W , is given by the Law of Gravitation with $r \approx R$ (radius of the Earth) as $W = \frac{GMm}{R^2}$, where M is the mass of the Earth. Weight is also given by Newton's Second Law. For a body falling under gravity with constant acceleration g close to the Earth's surface $W = mg$ and so $g \approx 9.81 \text{ m s}^{-2}$.

4. Forces (1)

12. Impulse & Momentum

Linear momentum, \underline{p} , of a body of mass, m , with velocity, \underline{v} , is a vector quantity defined as $\underline{p} = m\underline{v}$.

Impulse: If a constant force, \underline{F} , acts over a time, t , on the body then the impulse of the force is defined as $\text{Impulse} = \underline{F}t$. Impulse is a vector quantity. The unit of impulse is the same as the unit of momentum.

Relationship between momentum and impulse: If a force acts on a body over a time t , the impulse of the force equals the final momentum minus the initial momentum. For the case of a constant force,

$$\underline{F}t = m\underline{v} - m\underline{u}$$

Principle of conservation of linear momentum: When no resultant external force acts on a system of interacting (colliding) particles the total momentum of the system remains constant.

The collision of two bodies: An **elastic** collision is one in which the total kinetic energy is conserved. An **inelastic** collision is one in which the total kinetic energy always decreases. Consider the collision between two spheres moving in the same line.

Let

before collision after

m_1, m_2 = the masses of the two spheres
 u_1, u_2 = the velocities before collision
 v_1, v_2 = the velocities after collision
 $v_a = u_1 - u_2$ = the speed of approach
 $v_s = v_2 - v_1$ = the speed of separation

In a collision v_a and v_s are connected by the relation $v_s = e v_a$, or $v_2 - v_1 = e(u_1 - u_2)$ where $0 \leq e \leq 1$ and is called the **coefficient of restitution**.

In an elastic collision, $e = 1$. For an elastic collision $m_1 v_1 + m_2 v_2 = m_1 u_1 + m_2 u_2$

$$\frac{1}{2} m_1 v_1^2 + \frac{1}{2} m_2 v_2^2 = \frac{1}{2} m_1 u_1^2 + \frac{1}{2} m_2 u_2^2$$

In the case of spheres having the same mass ($m_1 = m_2$) $u_2 = v_1, u_1 = v_2$ which means the spheres exchange velocities.

In a 'perfectly inelastic' collision, where the bodies coalesce, $e = 0$. Then $v_1 = v_2$; there is no rebound, as shown.



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force, then the power is $P = Fv$. The rate at which work is done is called the **power**. If a constant force F is exerted on a body which moves with speed v in the direction of the force, then the power is $P = Fv$.

Power and Velocity: The rate at which work is done is and potential energies of the body, is conserved. total mechanical energy, which is the sum of the kinetic force acting on the body is the gravitational force, the **Conservation of Total Mechanical Energy:** When the only force acting on the body is the gravitational force, the total mechanical energy is conserved. So P.E. (gravitational) = mgh .

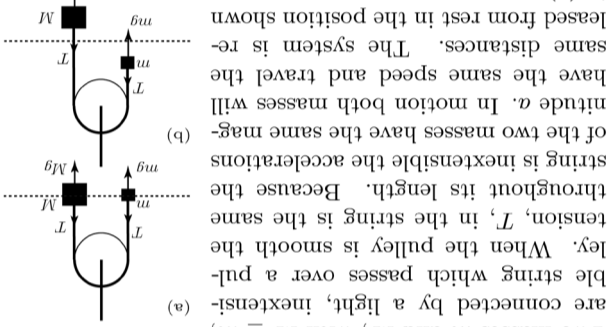
Potential Energy: P.E. is due to a body's position. mgh , of a body and the height, h , of its centre of gravity, **Gravitational Potential Energy** is the product of the weight, equal to the work done by the external forces on the body. K.E. = $\frac{1}{2}mv^2$. The change in the K.E. of a rigid body is body of mass m moves with speed v its K.E. is defined as **Kinetic Energy:** K.E. is due to a body's motion. When a gain or lose energy.

Energy: When a force does work on a body the body can placement the work done is zero. force is at right angles to the displacement. If the force is positive / negative respectively, the work done is positive / negative respectively, the work done is in the same / opposite direction as the displacement. If the component of the force is in the same direction as the displacement, work is a scalar quantity. $W = F \cdot s = F \cos \theta \cdot s$. Work is a scalar quantity.

Work done by a constant force: The figure below represents a body moving in a horizontal direction. A constant force, F , at an angle θ to the direction of motion, is exerted on the body. The work done, W , by the force, when its point of application undergoes a displacement s , is $W = F \cdot s = (F \cos \theta) \cdot s$. Work is a scalar quantity.

11. Work, Energy and Power

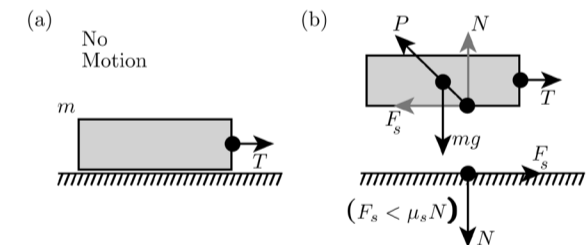
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10. Motion of connected particles

5. Forces (2)

Friction: The force which prevents, or tries to prevent, the slipping or sliding of two surfaces in contact is called **friction**. When the surface of one body slides over another, each body exerts a frictional force on the other, parallel to the surfaces. The frictional force on each body is opposite to the direction of its motion. Frictional forces may also act when there is no relative motion, as shown.



A cord is attached to a block of weight $W = mg$ and the tension, T , in the cord is such that the block remains at rest (diagram (a)). Diagram (b) is the corresponding separated body diagram. P is the force exerted on the block by the surface. N and F_s are the components of P , normal to and parallel to the surface. F_s is called the force of static friction. From Newton's 2nd law,

$$N = -W \quad \text{and} \quad F_s = -T$$

with corresponding scalar forms

$$N = W \quad \text{and} \quad F_s = T$$

As T is increased, a limiting value is reached after which the block starts to move. Thus there is a certain maximum value which F_s can have. The magnitude of this maximum value depends on the normal force N and a useful empirical law is

$$F_s(\text{max}) = \mu_s N$$

where μ_s is called the coefficient of static friction. The magnitude of the actual force of static friction can take any value between 0 and $F_s(\text{max})$. Thus

$$F_s \leq \mu_s N$$

As soon as sliding begins, the friction force decreases. This new friction force, F_k , also depends on the normal force. The empirical law used is

$$F_k = \mu_k N$$

where μ_k is the coefficient of sliding (or kinetic) friction. The values of μ_s and μ_k depend on the nature of the two surfaces which are in contact.

6. Kinematics: Rectilinear Motion

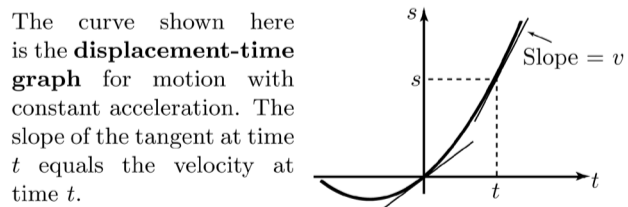
A **particle** is a body which can be modelled as a point mass in a given context. For example, for the motion of the planets about the Sun, then the Sun, Earth, etc., can be regarded as particles.

Kinematics is the study of the motions of particles and rigid bodies without any consideration of the forces required to produce these motions. Rectilinear motion is concerned with the motion of a single particle along a straight line.

Constant acceleration: The equations of motion are

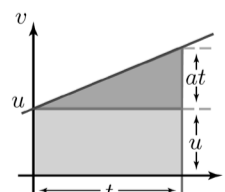
$$\begin{aligned} v &= u + at \\ s &= \frac{1}{2}(u + v)t \quad \text{or} \quad s = ut + \frac{1}{2}at^2 \\ v^2 &= u^2 + 2as \end{aligned}$$

where a is the (constant) acceleration, t represents time, v is the velocity at time t , u is the velocity at $t = 0$, s is the displacement at time t , and $s = 0$ at $t = 0$. These equations are obtained from $\frac{dv}{dt} = a$ and $\frac{ds}{dt} = v$.



The diagram here shows a **velocity-time graph** for rectilinear motion with constant acceleration. The area under a velocity-time graph equals the displacement. The gradient of the line represents the acceleration.

Non-constant acceleration: Here the acceleration, a , is a function of time, t . As for constant acceleration, the equations of motion are found by integrating $\frac{dv}{dt} = a(t)$ and $\frac{ds}{dt} = v$.



When the circular motion is uniform the speed, v , is constant and so the velocity and acceleration vectors are then

$$\vec{v} = v\hat{e}_\theta \quad \text{and} \quad \vec{a} = -\frac{v^2}{r}\hat{e}_r$$

So, if a particle of mass m moves uniformly in a circle of radius r , with speed v , the radial acceleration has magnitude v^2/r and is directed inward along the radius.

The conical pendulum: A particle of mass m revolves in a horizontal circle of length L . The cord makes an angle α with the vertical. The radius of the circle is $R = L \sin \alpha$. Hence $v = L \sin \alpha \omega$, where ω is the angular speed of motion in the horizontal circle. The forces exerted on the body are its weight, W , and the tension in the cord which resolves into horizontal and vertical components of magnitudes $T \sin \alpha$ and $T \cos \alpha$ resp. The body has no vertical acceleration and the radial acceleration has magnitude v^2/R . From Newton's 2nd law vertically and radially

$$T \cos \alpha - W = 0 \quad \text{and} \quad T \sin \alpha = \frac{mv^2}{R}$$

Then $\tan \alpha = \frac{v^2}{Rg}$ and $\cos \alpha = \frac{Rg}{\sqrt{R^2g^2 + v^4}}$. Motion arises only if $\cos \alpha < 1$, that is $\omega^2 > g/L$. If $\omega^2 > g/L$ then $\alpha = 0$.

Motion in a vertical circle: Consider a small body of mass m attached to a cord of length R and whirling in a vertical circle about O . The cord makes an angle θ , measured anti-clockwise from the downward vertical. The motion is circular but not uniform. The forces acting on the body are its weight, $W = mg$, and the tension T in the cord. The radial acceleration has magnitude v^2/R where $v = R\dot{\theta}$ (not necessarily constant).

From Newton's 2nd law in the radial direction the magnitude of the tension in the cord is $T = m\left(\frac{v^2}{R} + g \cos \theta\right)$. From Newton's 2nd law in the tangential direction $-mg \sin \theta = mR\ddot{\theta}$.

Writing θ as ω , $\frac{d^2\theta}{dt^2} = \frac{d\omega}{dt}$, this becomes $-mg \sin \theta = mR\omega \frac{d\omega}{d\theta}$ which by integrating gives the energy equation $\frac{1}{2}mR\omega^2 + mgR(1 - \cos \theta) = \frac{1}{2}mv^2$ where V is the speed when $\theta = 0$. The critical speed below which the cord becomes slack ($T = 0$) at its highest point (where $\theta = \pi$) is $v_c = \sqrt{Rg}$.

It follows that $v^2 = \omega^2 R^2 (A^2 - \cos^2 \theta)$. Here t is time, the amplitude, A , is the maximum value of $|x|$, v is the velocity, and ϵ is the initial phase angle. The period, τ , is the time for a complete oscillation. The frequency, f , is the number of oscillations per unit time. ω is the angular frequency given by $\omega = 2\pi f = \frac{2\pi}{\tau} = \sqrt{\frac{k}{m}}$. The graph shows $x(t)$ for the case $\epsilon = 0$. The initial position of the particle is its maximum positive displacement. The maximum speed occurs when $x = 0$, i.e. at the centre of the oscillation. The acceleration is maximum when the displacement x is maximum.

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